



# History and Significance of the “Spin-Statistics Connection” (the subject of our conference)

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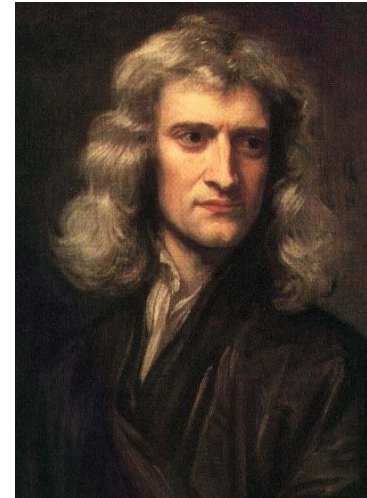
# Outline

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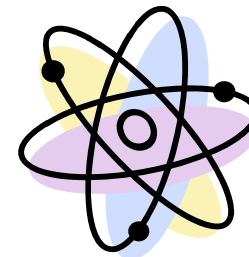
- Basic science issues :
  - Identical particles
  - **Fermions** and **bosons**
  - Behavior of groups of fermions and bosons  
= "**statistics**"
- Some history of these issues
- Important consequences
- Take-home messages

# Identical Particles – same mass, electrical charge, ....

- **Newtonian Physics** – identity is of no consequence



- **Quantum Physics** – very important!



# Identical Particles

- **Example:** all **electrons** have the same mass, electrical charge, magnetic properties... Therefore, we cannot distinguish one electron from another.
- Quantum theory then restricts the **kinds of states** for electrons
- Both theoretical and practical importance (**Example:** electrons in **semiconductors**).

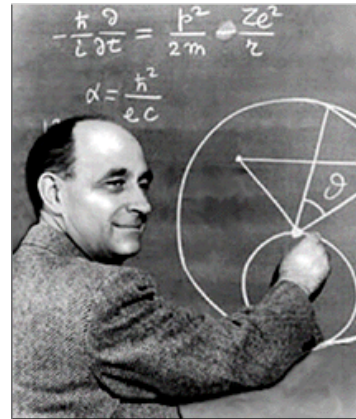
Watch the iPod touch tv ad.



# Two Families of Fundamental Particles

## ☒ Fermions

Enrico Fermi  
1901-1954



## ❖ Bosons

Satyendranath Bose  
1894-1974



# Fermions

- **electrons, neutrinos, protons, quarks...**
- **At most one particle per quantum state**

"statistics"





# Bosons

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➤ photons, W, Z bosons, gluons,  
 $^4\text{He}$  (nucleus and atom),  $^{16}\text{O}$   
(nucleus),  $^{85}\text{Rb}$  (atom),...

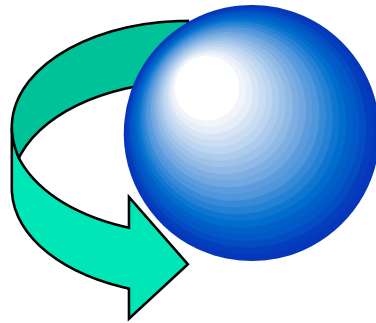
➤ Possibly many particles in the  
same quantum state

“statistics”



# What distinguishes Fermions from Bosons?

➤ Spin (angular momentum)







# Fermions versus Bosons

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- **Quantum theory** tells us that angular momentum (amount of spin) =  $n (h/2)$  ←

Experimental fact:

- **Fermions** have  $n = 1, 3, 5, \dots$
- **Bosons** have  $n = 0, 2, 4, \dots$

# Fermions versus Bosons?

- Why spin?
- Richard Feynman  
1918-1988



Library of Congress (AP photo)



# Feynman quote

(Feynman Lectures Vol. III, Chapter 4)

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An explanation has been worked out by (Wolfgang) Pauli from complicated arguments of **Quantum Field Theory** and **relativity**...but we haven't found a way of reproducing his arguments on an elementary level...this probably means that we **do not have a complete understanding** of the fundamental principle involved...



# Outline of the History

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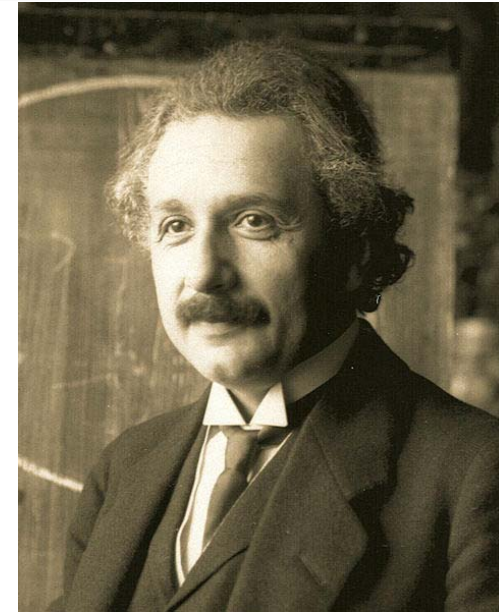
- 1924-25 **Exclusion Principle** for electrons, **Bose-Einstein** behavior for photons (1920 Bose, 1924 Einstein)
- 1925 **Spin**: intrinsic angular momentum
- 1926 **Identical Particles** in Quantum Theory
- 1926-29 Nuclei have **spin** angular momentum
- 1929-30 **The Connection**: established through molecular spectroscopy

# Photons are Bosons

## 1924



Satyendranath Bose  
1894 - 1974



Albert Einstein  
1879-1955

Many photons can be in the **same quantum state**.  
But they said nothing about spin angular momentum.

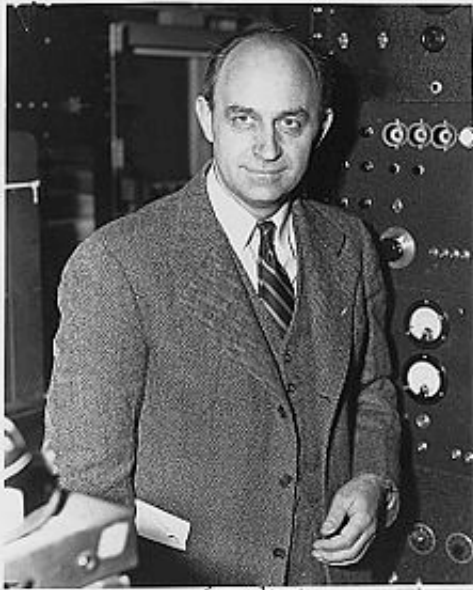
# Wolfgang Pauli



1925 Exclusion Principle explains atomic structure

**principio di esclusione di Pauli**

# Pauli-Fermi Principle



Two **electrons** cannot  
be in the same  
quantum state.

Today, we would say "**Electrons are fermions.**"

# Electron Spin



Ralph Kronig  
1904-1995

- Kronig was first to suggest that electrons have **spin angular momentum**.

**Pauli** especially ridiculed the idea of spin, saying that "it is indeed very clever but of course has nothing to do with reality."



# Heisenberg and Dirac

## 1926-27 – Identical Particles



Werner Heisenberg – 1901-1976 Paul Dirac – 1902-1984

Quantum states with **identical particles** are either **symmetric** or **anti-symmetric** if the “coordinates” of the particles are interchanged.

# Identical Particles in Quantum Theory



- Symmetric states - bosons
- Anti-symmetric states - fermions

Quantum theory restricts the kinds of states for identical particles.

Anti-symmetric states: at most one particle per state.

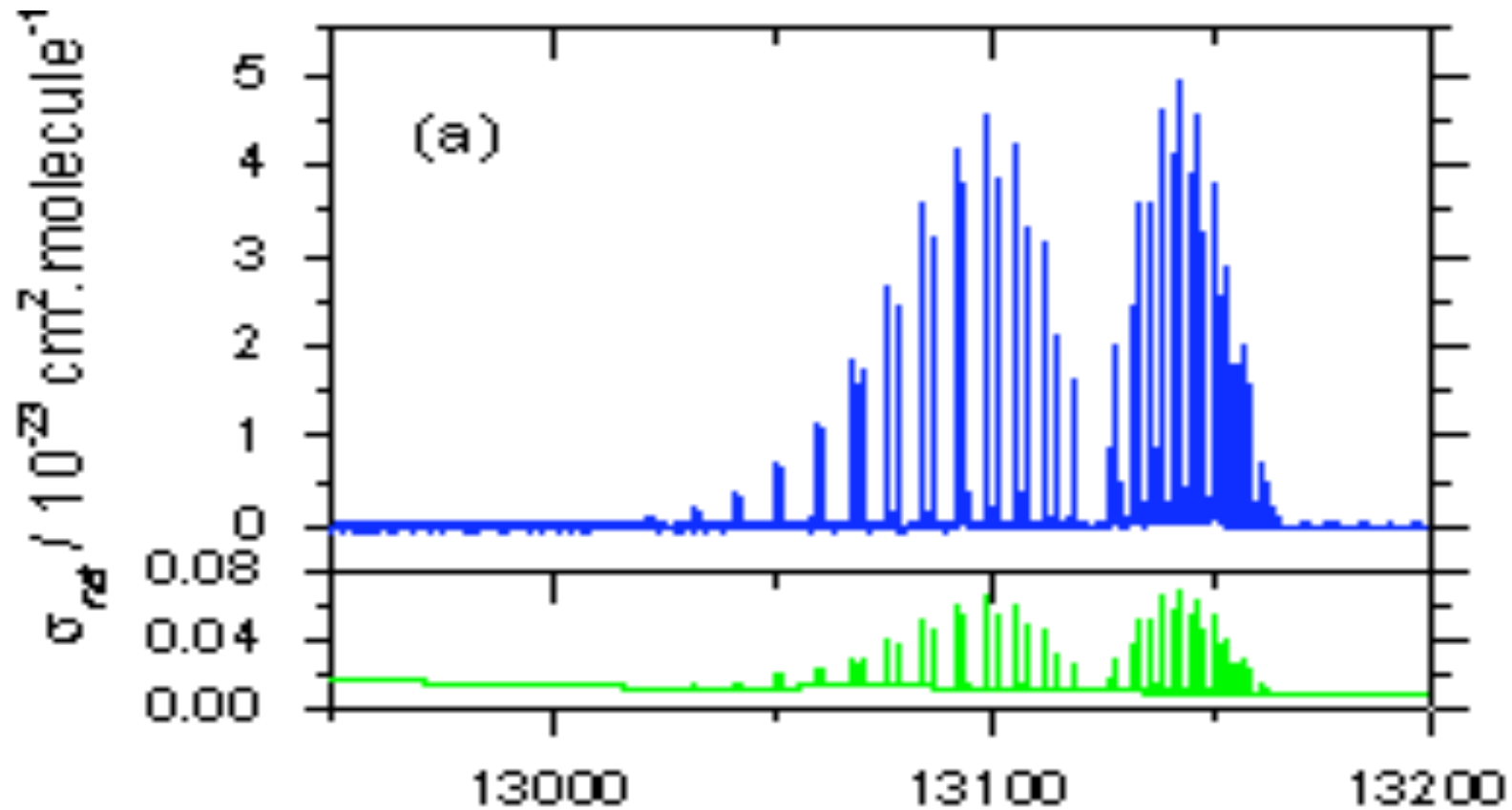


# Molecular Spectroscopy

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- 1925 Some molecules show **alternating intensities** (or **missing lines**) in rotationally resolved spectra
- 1925 **R. Mecke**: these are always **homo-nuclear** molecules (identical nuclei)
- 1929-30 **F. Rasetti**, Raman spectroscopy of  $\text{H}_2$ ,  $\text{N}_2$ ,  $\text{O}_2$

# O<sub>2</sub> Spectrum near 762 nm



# Fredrich Hund (1896-1997)



1927 – molecular spectroscopy – Hund associated alternating intensities with the **spin angular momentum of nuclei** in the molecules



# Franco Rasetti (1901-2001)

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- In Fermi's group in Roma
- In 1929 went to the [California Institute of Technology](#) to work with Robert A. Millikan
- Carried out Raman spectroscopy at CalTech (a newly discovered technique)

# Franco Rasetti



From his Raman spectroscopy on  $H_2$ ,  $O_2$ , and  $N_2$ :

He recognized the connection between the **alternating intensities** and the **spin angular momentum** of the nuclei in the diatomic molecule and **the symmetry** and **anti-symmetry** of the states.



# The facts – a review

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- Quantum theory tells us that angular momentum (amount of spin) =  $n (h/2)$
- **Fermions** have  $n = 1, 3, 5, \dots$   
(anti-symmetric states)
- **Bosons** have  $n = 0, 2, 4, \dots$   
(symmetric states)



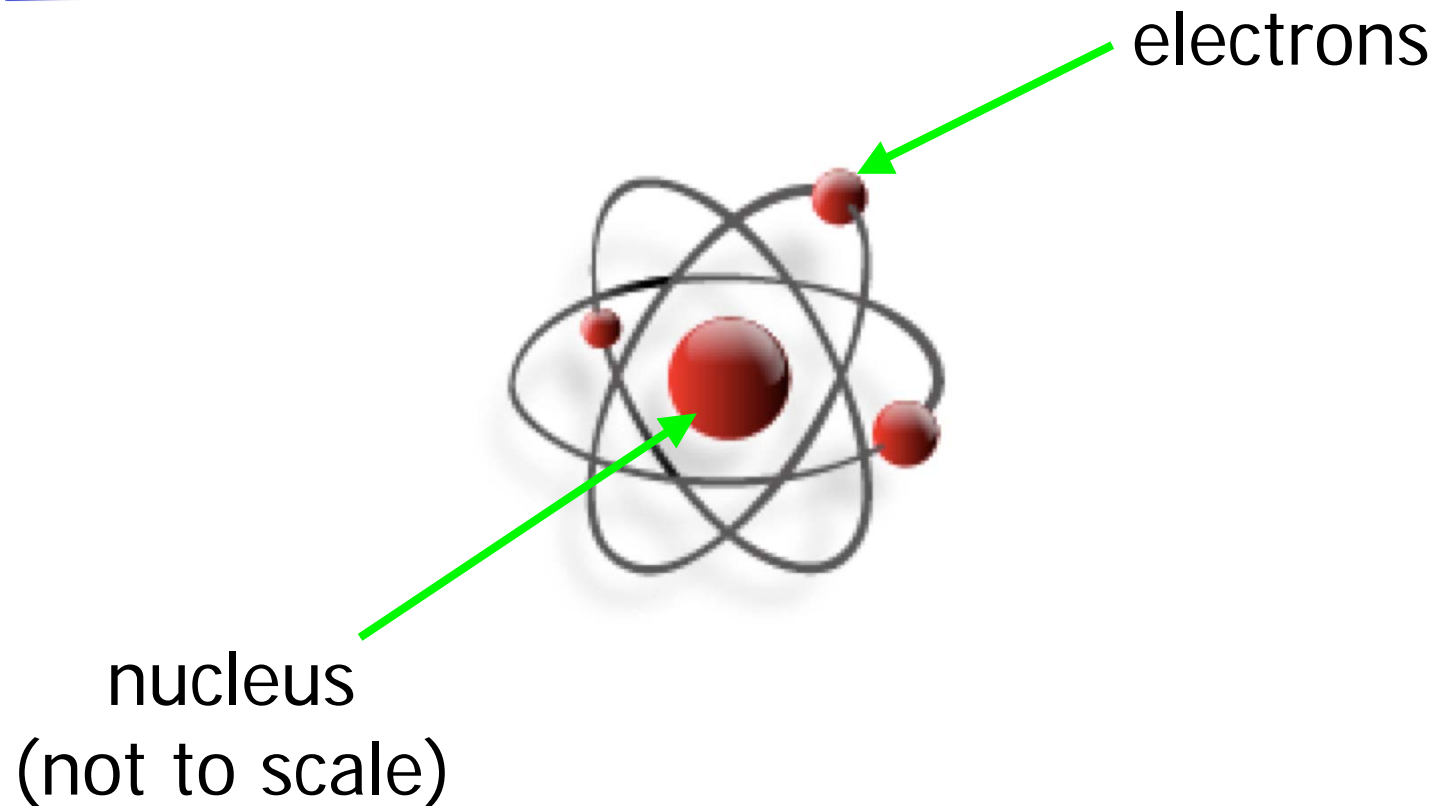


## But why spin?

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- Pauli 1940 – “proof” using **Quantum Field Theory** (quantum theory plus Einstein’s theory of relativity).
- Many other scientists.....
- .....
- This conference.....

# Consequences: I. Atomic structure and chemistry



# Periodic Table of the Elements

**Periodic Table of the Elements**

1 H																	2 He														
3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne														
11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar														
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr														
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe														
55 Cs	56 Ba	57 La	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn														
87 Fr	88 Ra	89 Ac	104 Unq	105 Unp	106 Unh	107 Uns	108 Uno	109 Une	110 Unn																						
																		58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu
																		90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr

- hydrogen
- alkali metals
- alkali earth metals
- transition metals
- poor metals
- nonmetals
- noble gases
- rare earth metals

# Consequences:

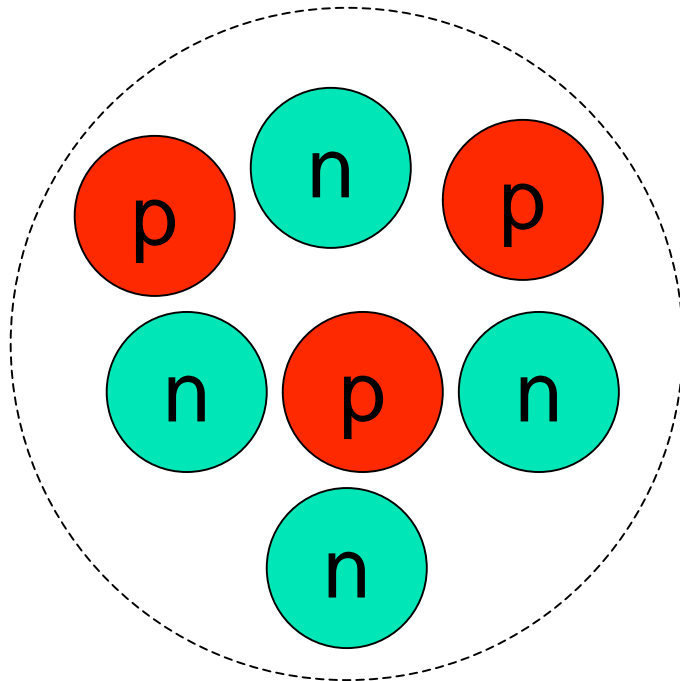
## II. Stability of matter



- Matter doesn't collapse – **The Pauli Exclusion Principle** is important to understand why.

Consequences:

### III. Stability of atomic nucleus

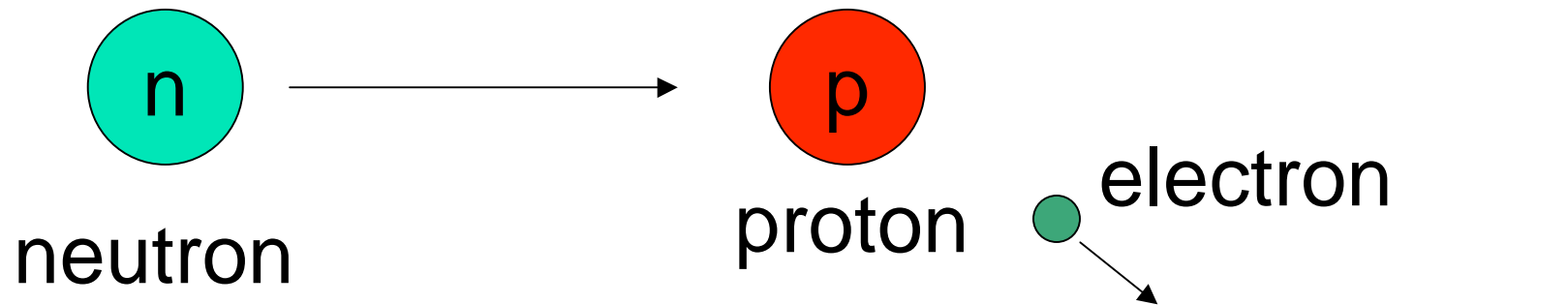


${}^7\text{Li}$  nucleus

3 protons

4 neutrons

# Outside a nucleus: Neutrons are unstable.

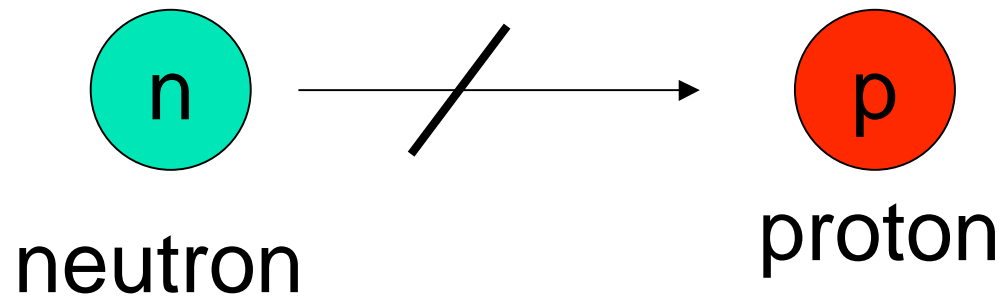


“Half life” about 10 minutes



# Inside a nucleus

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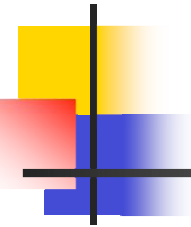


Because all of the accessible  
proton states are **already occupied**.

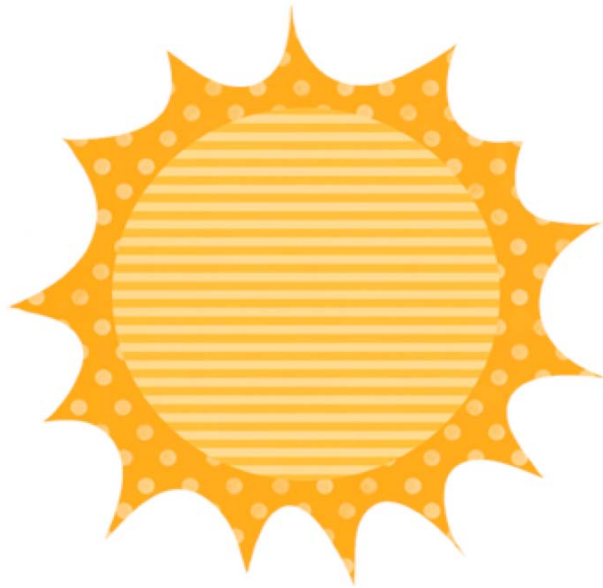
The Pauli Exclusion Principle!

# Consequences:

## IV. Neutron stars



star

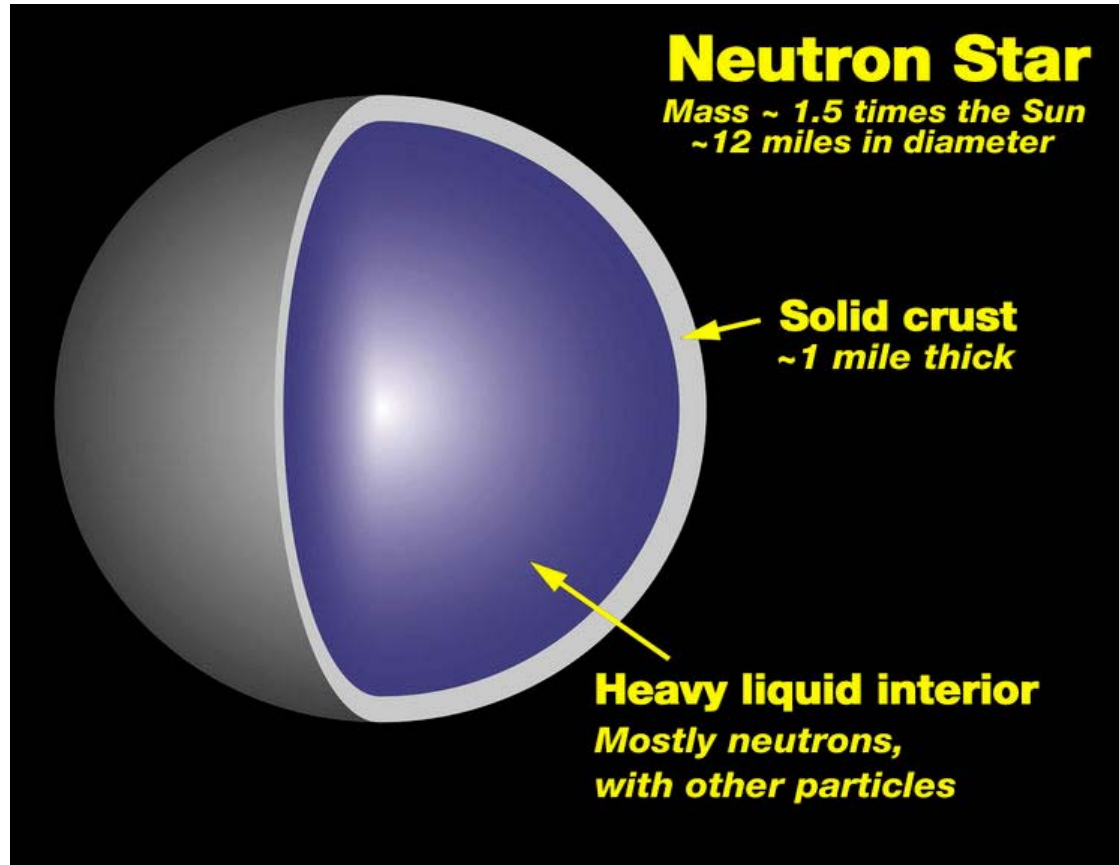
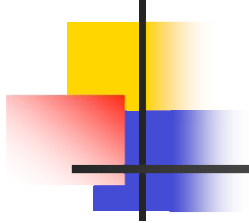


supernova

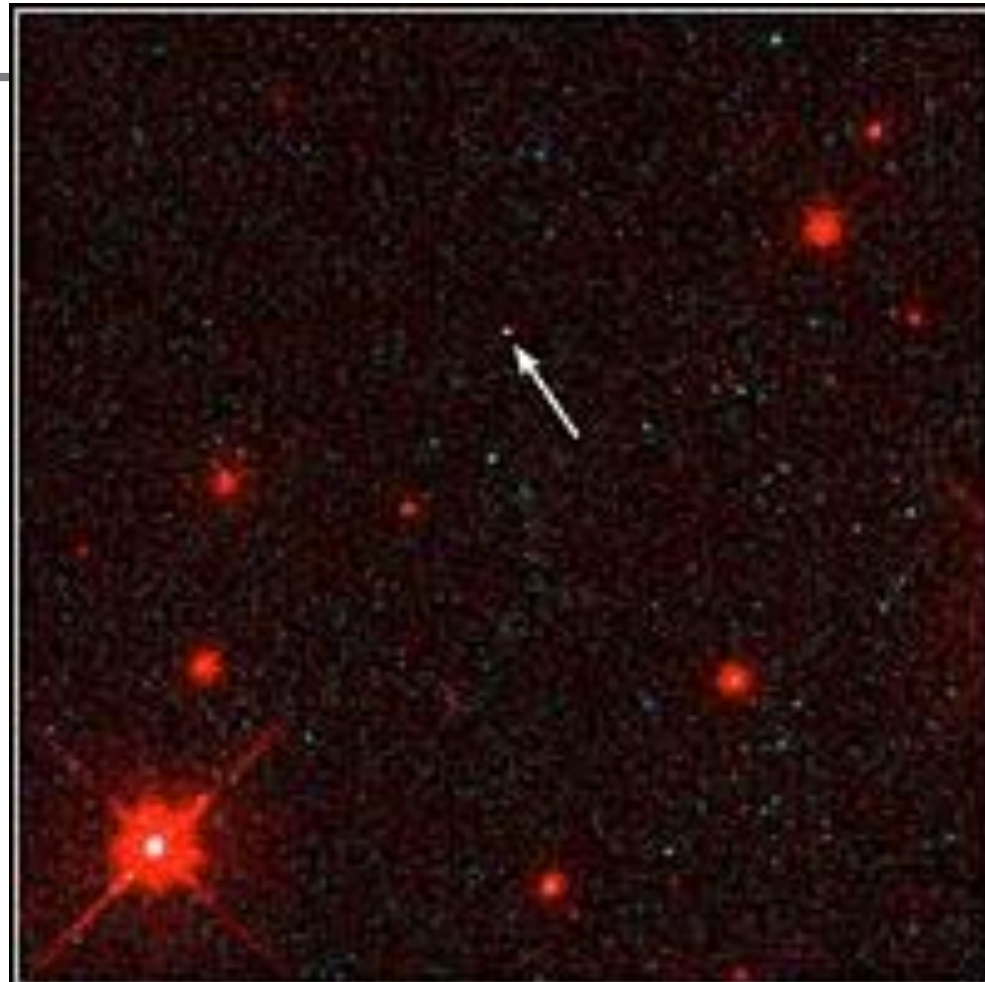


Neutron star in the  
Crab Nebula  
1054 AD





# Visible Neutron Star



**Isolated Neutron Star RX J185635-3754** HST • WFPC2  
PRC97-02 • ST ScI OPO • September 25, 1997  
F. Walter (State University of New York at Stony Brook) and NASA



# Conclusions

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- Without the **Pauli Exclusion Principle**, the world would be **very different** (and we might not be here).
- We should try to understand **where the Pauli Exclusion Principle comes from** and its limitations.



# Take-home Messages

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- Even very intelligent scientists can be **confused for many years** over fundamental principles.
- So, **don't give up** if you don't understand something immediately.  
**Keep asking questions!**

Grazie e Buona Sera!

